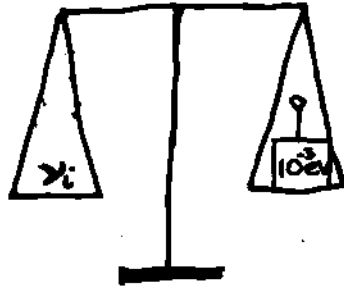


Neutrino Mass
and
Extra Dimensions

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Introduction

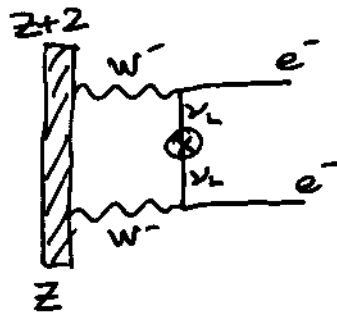
Neutrinos have mass :



But, are neutrinos Majorana or Dirac in nature ?

Experiment Look for lepton number violation :

e.g. neutrinoless double β decay ($0\nu\beta\beta$)



Theory Seesaw mechanism is elegant and simple, but can also have natural Dirac mass!

① Majorana

At low energies (IR) lepton number is violated by an irrelevant op. (dim. 5) :

$E < M$ $\mathcal{L} = \mathcal{L}_{SM} + \frac{\lambda_M}{M} \nu_L \nu_L H H$

$\langle H \rangle \sim v \Rightarrow m_\nu = \lambda_M \frac{v^2}{M} = \lambda_M v \cdot \frac{v}{M}$ extra suppression

\Rightarrow naturally small provided $M \gg v$

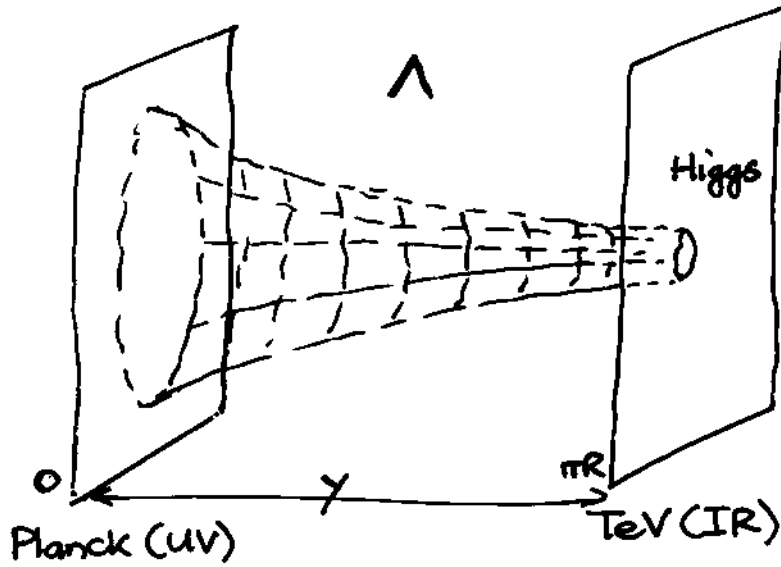
$E = M$ Introduce RH neutrino ν_R with mass term

$\mathcal{L} = M \nu_R \nu_R + h.c.$

See saw mechanism

$(\nu_L \nu_R) \begin{pmatrix} 0 & m \\ m & M \end{pmatrix} \begin{pmatrix} \nu_L \\ \nu_R \end{pmatrix} \Rightarrow m_\nu = \frac{m^2}{M}$, M
~
RH neutrino
decouple for
 $E < M$

Warped Extra Dimensions [Randall, Sundrum (1999)]

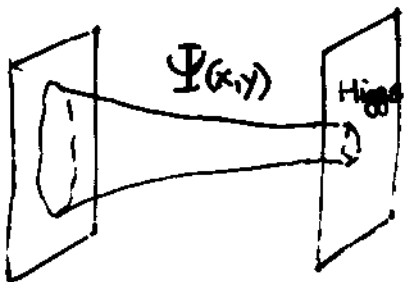


Metric: $ds^2 = e^{-2ky} dx^2 + dy^2$ \rightarrow slice of AdS_5 ($\Lambda < 0$)

warp factor

where $k = AdS$ curvature scale ($k \sim M_p$)

Fermions in the bulk [TG, Pomarol (2000)]



$$\Psi_i(x,y) = \begin{pmatrix} \psi_{1i} \\ \bar{\psi}_{2i} \end{pmatrix} \Rightarrow \begin{array}{c} \vdots \\ \text{orbifold} \\ \begin{array}{l} (3) \text{---} \\ (2) \text{---} \\ (1) \text{---} \\ (0) \text{---} \end{array} \\ \psi_{1i} \quad \bar{\psi}_{2i} \end{array}$$

$$S_{\text{bulk}} = -i \int d^4x dy \sqrt{-g} \left[\bar{\Psi}_i \Gamma^M D_M \Psi_i + \underbrace{c_i k \epsilon(y)}_{\text{bulk mass term}} \bar{\Psi}_i \Psi_i \right]$$

Zero mode $\psi_{1i}^{(0)}(y) \sim \frac{1}{N_0} e^{(\frac{1}{2} - c_i)ky}$

Cases (i) $c_i > \frac{1}{2} \Rightarrow$ localised towards Planck brane

(ii) $c_i < \frac{1}{2} \Rightarrow$ " " TeV brane

② Dirac

Introduce RH neutrino that does not decouple

$$\mathcal{L} = \mathcal{L}_{SM} + \lambda_D \nu_L \nu_R H$$

$$\langle H \rangle \sim v \Rightarrow m_\nu = \lambda_D \underbrace{v}_{\text{no suppression!}}$$

Requires: (i) $\Delta m_\nu^2 \sim 10^{-4} \text{eV}^2 \Rightarrow \lambda_D \sim 10^{-13}!$

(ii) No dim 3 relevant operator: $M \nu_R \nu_R$

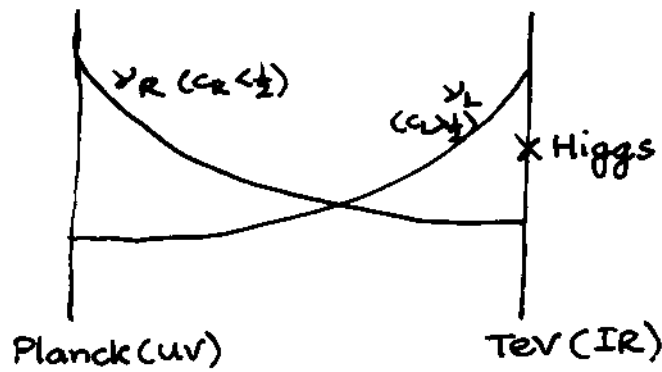
→ Impose global lepton number symmetry

But, badly violated in the UV → quantum gravity
⇒ $M = M_P!$

How can (i) and (ii) be natural?

→ use warped extra dimensions!

Dirac neutrino mass [Grossman, Neubert (2000)]

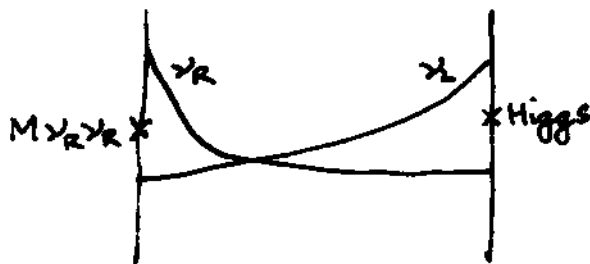


Tiny wavefunction overlap $\Rightarrow \lambda_0 \ll 1!$

But, at Planck brane expect global lepton number to be violated

\Rightarrow neutrino no longer Dirac!

In fact [Huber, Shafi (2003)]



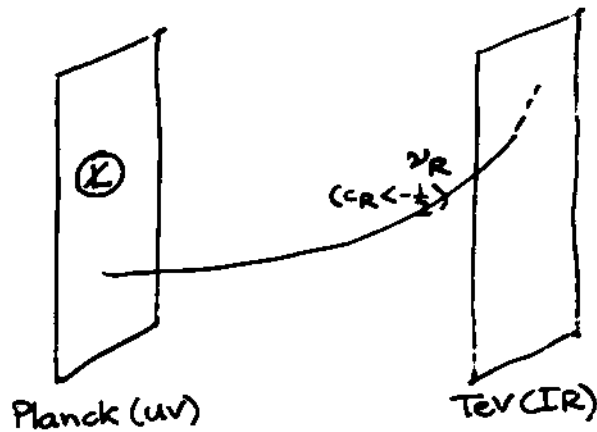
$\Rightarrow m_\nu \sim \frac{v^2}{M}, M$

i.e. usual seesaw mechanism

But, can we still have a Dirac neutrino?

Yes!

Consider RH neutrino localised on TeV brane [T.G. (2004)]

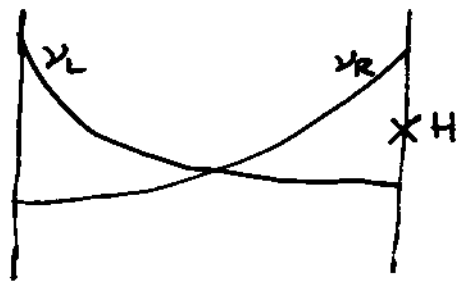


$$S_{\text{total}} = S_{\text{bulk}} + \underbrace{\int d^4x \int dy \, i \frac{b_M}{2} \delta(y) (\bar{\Psi}_R^c \Psi_R + \text{h.c.})}_{\text{boundary Majorana mass}}$$

$$\begin{aligned} \text{Majorana mass: } m_0 &\sim \frac{b_M}{2} (1-2c_R) k e^{-(1-2c_R)\pi k R} \\ &\sim \frac{b_M}{2} \left(\frac{\text{TeV}}{M_p}\right)^{1-2c_R} \cdot M_p \quad \rightarrow 0 \text{ as } c_R \rightarrow -\infty \\ &\quad \underbrace{\hspace{1cm}}_{\text{huge suppression for } c_R < -\frac{1}{2}} \end{aligned}$$

\Rightarrow RH neutrino does not decouple

But we still need to generate a mass:



Consider Yukawa interaction:

$$S_Y = \int d^4x \int dy \underbrace{\lambda_5}_{\text{5D Yukawa coupling}} \nu_L(x,y) \nu_R(x,y) H(x) \delta(y - \pi R)$$

$$\Rightarrow \lambda_4 \approx \lambda_5 k \sqrt{(c_L - \frac{1}{2})(\frac{1}{2} - c_R)} e^{(\frac{1}{2} - c_L)\pi k R} \quad (c_L > \frac{1}{2}, c_R < \frac{1}{2})$$

$$\Rightarrow \lambda_4 \approx 10^{-13} \text{ for } c_L \sim 1.36$$

Actually, to obtain realistic fermion mass spectrum \Rightarrow delocalise Higgs

$$\text{e.g. } c_H = 0.5; c_L = 1.32; c_{eR} = 0.2; c_{\nu R} = -2$$

$$\text{or } c_H = 0.11; c_L = 1.32; c_{eR} = 1; c_{\nu R} \approx -2.5$$

Thus, have

natural Dirac mass in presence of Planck scale lepton number violation!

Dual 4D interpretation

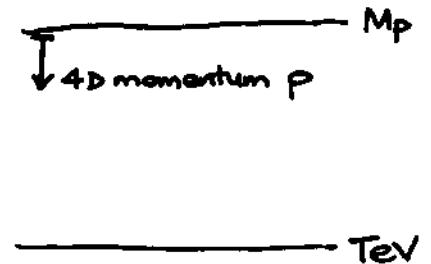
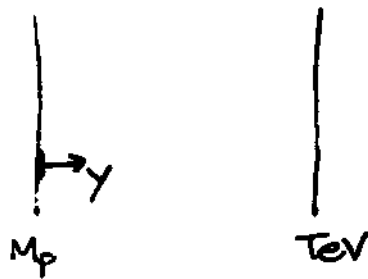
[Maldacena (1998)
Arkani Hamed, Porrati, Randall (2001)
Rattazzi, Zaffaroni (2001)]

Using AdS/CFT correspondence :

5D warped
extra dimension

DUAL

4D strongly coupled
CFT at large N



↔

UV bdy value
of bulk field
 $\Phi(x, y=0) \equiv \Phi(x)$

↔

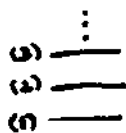
source of CFT
operators \mathcal{O}
 $\mathcal{L} = \lambda \Phi(x) \mathcal{O}(x)$

Bulk mass
 m_Φ

↔

dim \mathcal{O}

Kaluza-Klein
tower



↔

Resonances of
large N gauge theor.

Zero mode
near Planck brane

↔

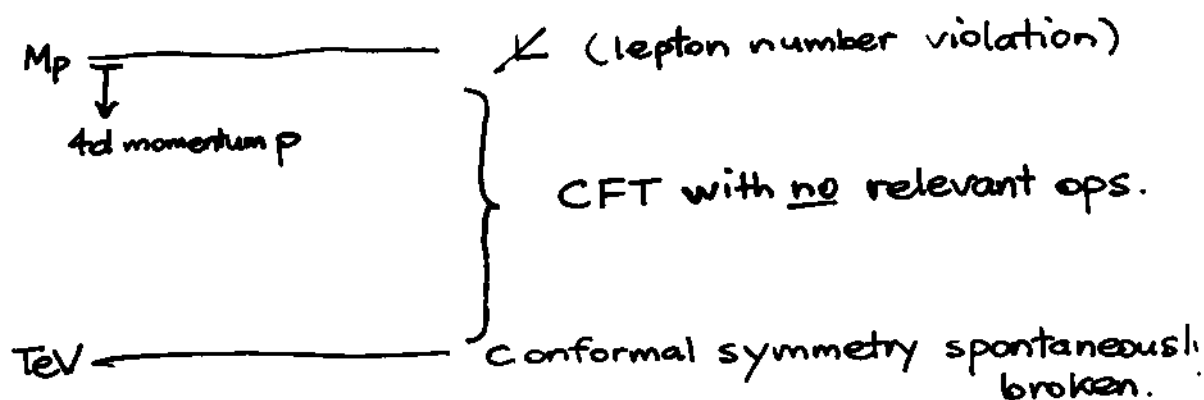
elementary state
added to CFT

Zero mode
near TeV brane

↔

bound state of
CFT

Dual interpretation of Dirac neutrino [TG (2004)]



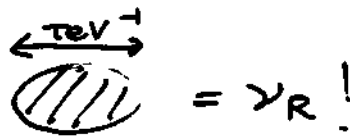
RH neutrino dual ^{4D} Lagrangian:

$$\mathcal{L} = \mathcal{L}_{\text{CFT}} + \frac{1}{\Lambda} k^{\frac{1}{2} + c_R} \underbrace{\bar{\Psi}_R}_{\text{elem. field}} \underbrace{\mathcal{O}_R}_{\text{CFT op.}} + \text{h.c.} \quad (c_R < -\frac{1}{2})$$

where $\dim \mathcal{O}_R = \frac{3}{2} + |c_R - \frac{1}{2}|$

In limit $c_R \rightarrow -\infty$, mixing term $\rightarrow 0$ (irrelevant)

\Rightarrow RH neutrino is a CFT bound state of size TeV^{-1} !



Also: Arkani-Hamed & Grossman (1998)

\rightarrow Not sensitive to UV violation of lepton number

$$\nu_R = \psi_{\mathcal{O}_R}^{(0)} - b_M \left(\frac{\mu}{M_p}\right)^{\frac{1}{2} - c_R} \psi_R$$

Decouples as $c_R \rightarrow -\infty$

$\Rightarrow \mathcal{L}_{\text{Majorana}} = b_M k \left(\frac{\mu}{R}\right)^{\frac{1}{2} - c_R} \nu_R \nu_R$

Dimension 3 mass term becomes highly irrelevant in IR!

LH neutrino dual^{4D} Lagrangian:

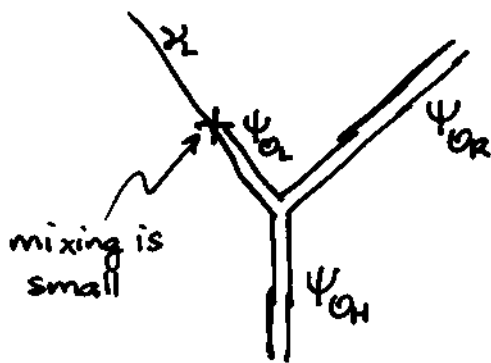
$$\mathcal{L} = \mathcal{L}_{\text{CFT}} + i \bar{\Psi}_L \bar{\sigma} \cdot \partial \Psi_L + \xi R^{\frac{1}{2} - c_L} \underbrace{\bar{\Psi}_L}_{\text{elem. field}} \underbrace{\mathcal{O}_L}_{\text{CFT op.}} + \text{h.c.} \quad (c_L)$$

where $\dim \mathcal{O}_L = \frac{3}{2} + |c_L + \frac{1}{2}|$

In limit $c_L \rightarrow +\infty$, mixing term $\rightarrow 0$

\Rightarrow LH neutrino is an elementary state

Dirac mass:



$$\mathcal{L}_{\text{Yukawa}} = \left(\frac{M}{R}\right)^{c_L - \frac{1}{2}} \nu_L \nu_R H +$$

\Rightarrow tiny Yukawa coupling

Conclusion

Composite
RH neutrino, ν_R
at TeV scale

Planck scale
lepton number
violation



Natural Dirac mass

Open questions

- Dirac leptogenesis
→ "Neutrino genesis" Dirk, Lindner, Ratz, Wright (2000)
- How does compositeness affect early universe cosmology?