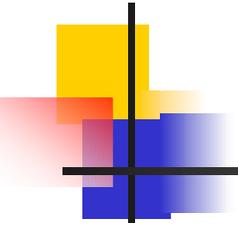


# Sensitivity of OPERA to $\theta_{13}$

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M. Komatsu, P.M., F. Terranova

A note with more details is available as [hep-ph/0210043](https://arxiv.org/abs/hep-ph/0210043)



# Motivations

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- **Scientific motivation**

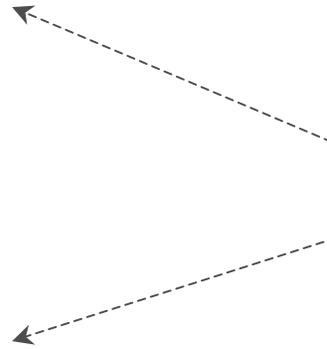
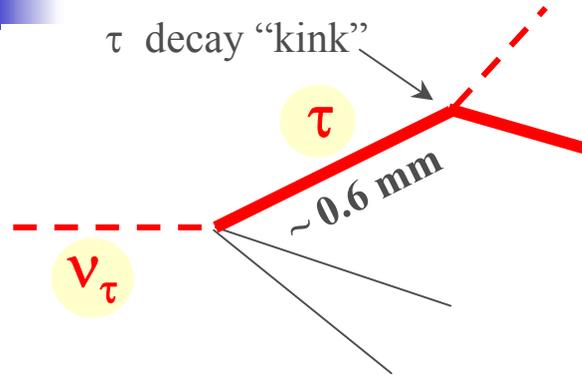
- the search for  $\theta_{13}$  is a very hot topic

- **Non-scientific motivations**

- the common wisdom among neutrino physicist is that OPERA has no sensitivity to  $\theta_{13}$
- in the paper V.D. Barger et al., Phys.Rev.D65(2002)053016 it is claimed that OPERA does not contribute to the  $\theta_{13}$  sensitivity of next LBL experiments
- different assumptions are made for different experiments running on the same beam

The  $\theta_{13}$  search in OPERA is performed in parallel with the  $\nu_{\tau}$  appearance search

# To identify $\tau$ leptons, "see" their decay topology



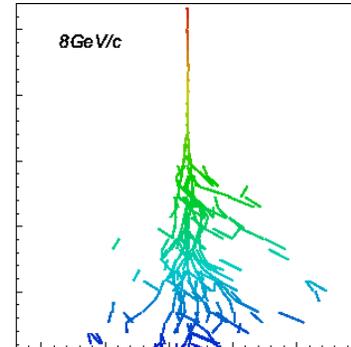
## The challenge

$\nu$  oscillation  $\rightarrow$  massive target **AND** decay topology  $\rightarrow$  micron resolution

Lead - nuclear emulsion sandwich  
"Emulsion Cloud Chamber", in brief "ECC"

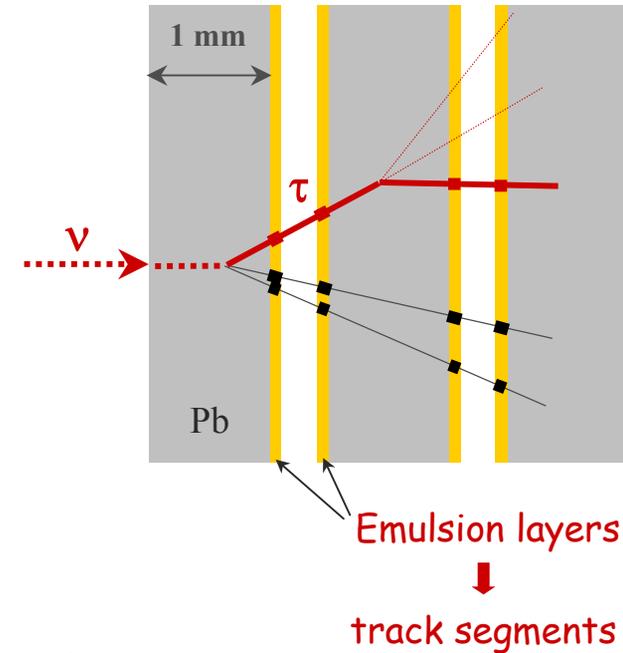
**BUT the ECC also provides:**

- electron and gamma detection plus energy measurement
- momentum measurement by multiple scattering



# The Emulsion Cloud Chamber (ECC)

- Emulsions for tracking, passive material as target
  - ↳  $< \mu\text{m}$  space res.
  - ↳ mass
- Established technique
  - charmed "X-particle" first observed in cosmic rays (1971)
  - DONUT/FNAL beam-dump experiment:  $\nu_\tau$  observed (2000)



$$\Delta m^2 = \mathcal{O}(10^{-3} \text{ eV}^2) \rightarrow M_{\text{target}} \sim 2 \text{ kton}$$

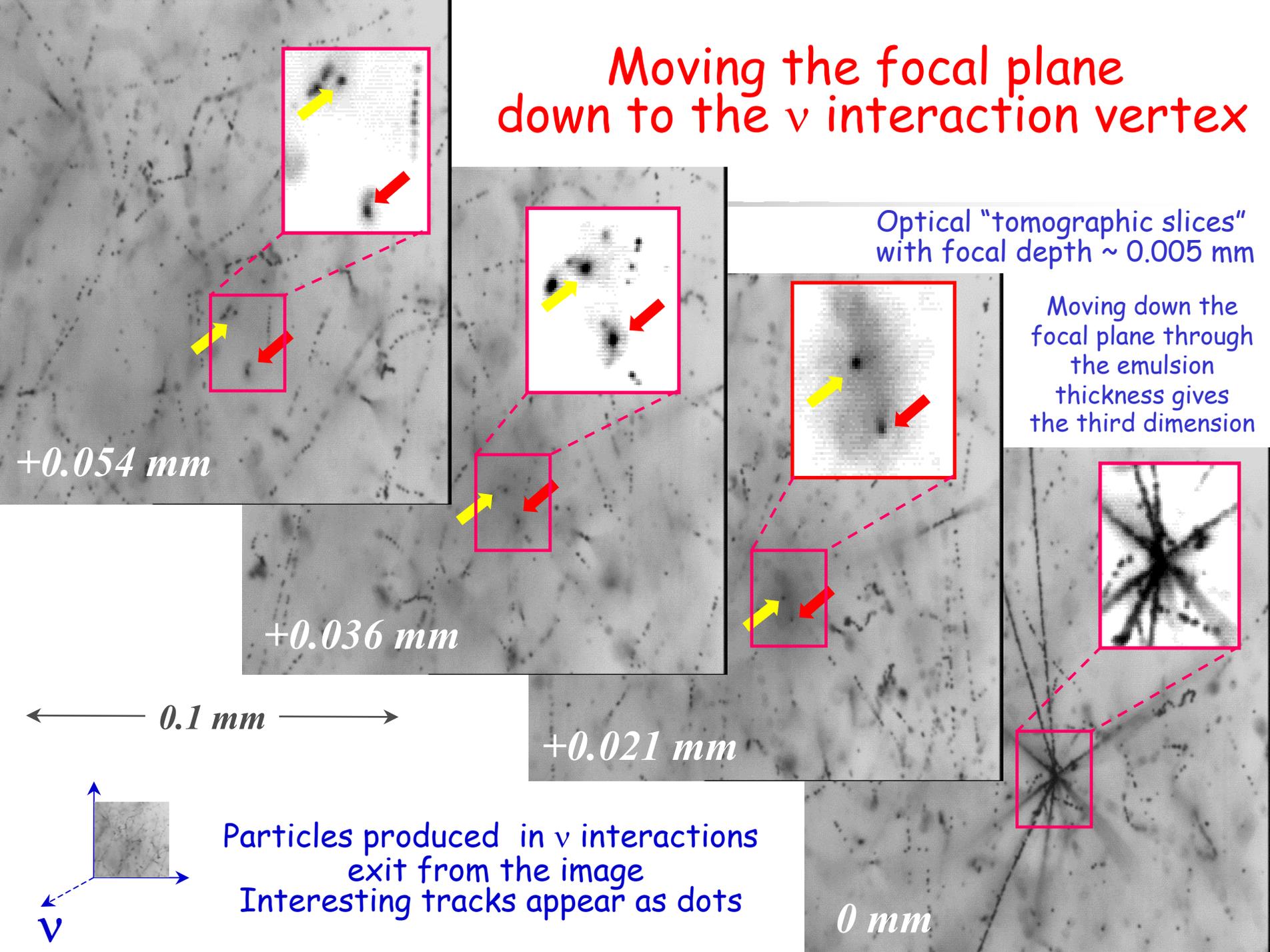
modular structure ("bricks"): basic performance is preserved  
large detector  $\rightarrow$  sensitivity, complexity  
required: "industrial" emulsions, fast automatic scanning

Experience with emulsions and/or  $\nu_\tau$  searches : E531, CHORUS, NOMAD and DONUT

# Moving the focal plane down to the $\nu$ interaction vertex

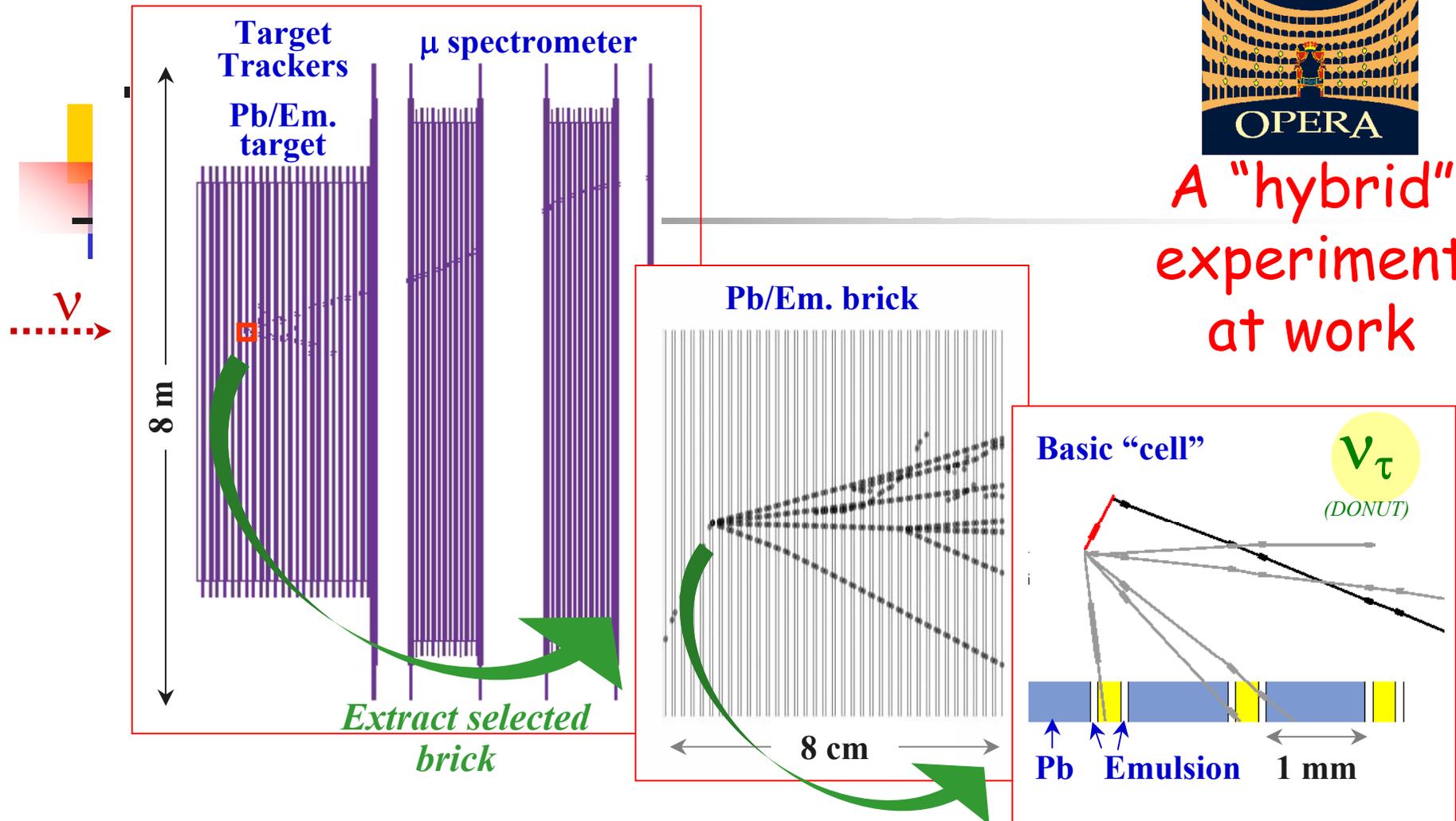
Optical "tomographic slices" with focal depth  $\sim 0.005$  mm

Moving down the focal plane through the emulsion thickness gives the third dimension





A "hybrid"  
experiment  
at work



### Electronic detectors

- select  $\nu$  interaction brick
- $\mu$  ID, charge and  $p$

### Emulsion analysis

- vertex search
- decay search
- $e/\gamma$  ID, kinematics

# The CNGS and the expected number of events in OPERA

Nominal  $\nu$  beam (Nov. 2000)

Shared SPS operation

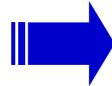
200 days/year

$4.5 \times 10^{19}$  pot / year

5 year run

1.65 kton average target mass

(accounting for mass reduction with time, due to brick removal for analysis)



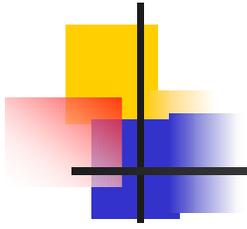
Expected interactions

$\sim 30000 \nu_{\mu}$  NC+CC

$\sim 120 \nu_{\tau}$  CC

at  $\Delta m^2 = 2.5 \times 10^{-3} \text{ eV}^2$  and full mixing

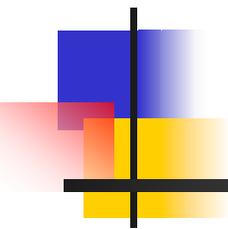
Limiting factor for  $\theta_{13}$  search:  
 $\nu_e + \text{anti-}\nu_e$  beam contamination  $\sim 0.87\%$



# Beam systematics

---

- The main systematic uncertainty comes from K production
- We assume for consistency with other works a 5% error on the  $\nu_e$  flux
- Given the small number of expected events in OPERA (see later) the sensitivity to  $\theta_{13}$  is dominated by the statistical fluctuations of the background
  - A systematic error up to 10% does not change appreciably the experimental sensitivity
- With the OPERA detector it is possible to (thanks to the spectrometer)
  - Measure the  $\mu^-$  energy spectrum (at high-energy  $\nu_\mu$  from  $K^+$  decays dominate)
  - Measure the  $\mu^+$  energy spectrum (anti- $\nu_\mu$  from  $K^-$  decays dominate)
    - ➡ Good samples ( $O(1\text{Kevts})$ ) to cross-check the beam simulation

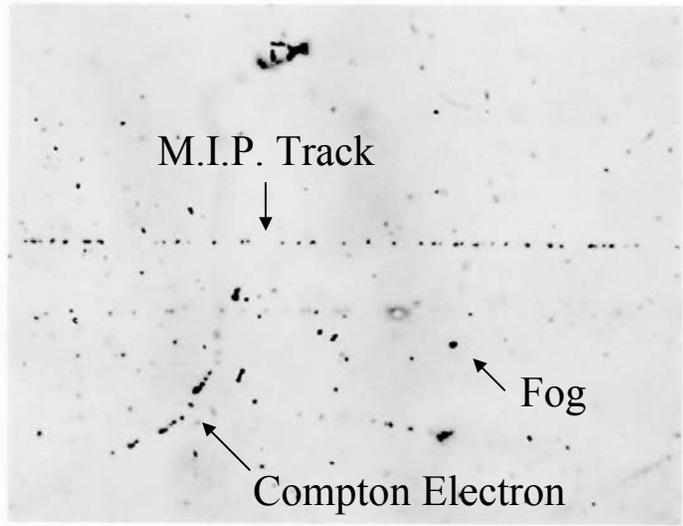


# An overview of possible measurements with emulsions

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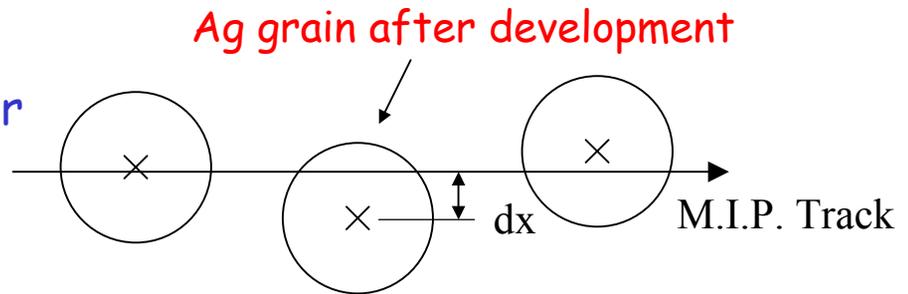
# Intrinsic space resolution in tracking with emulsion

Cross sectional view of an emulsion layer



100  $\mu\text{m}$

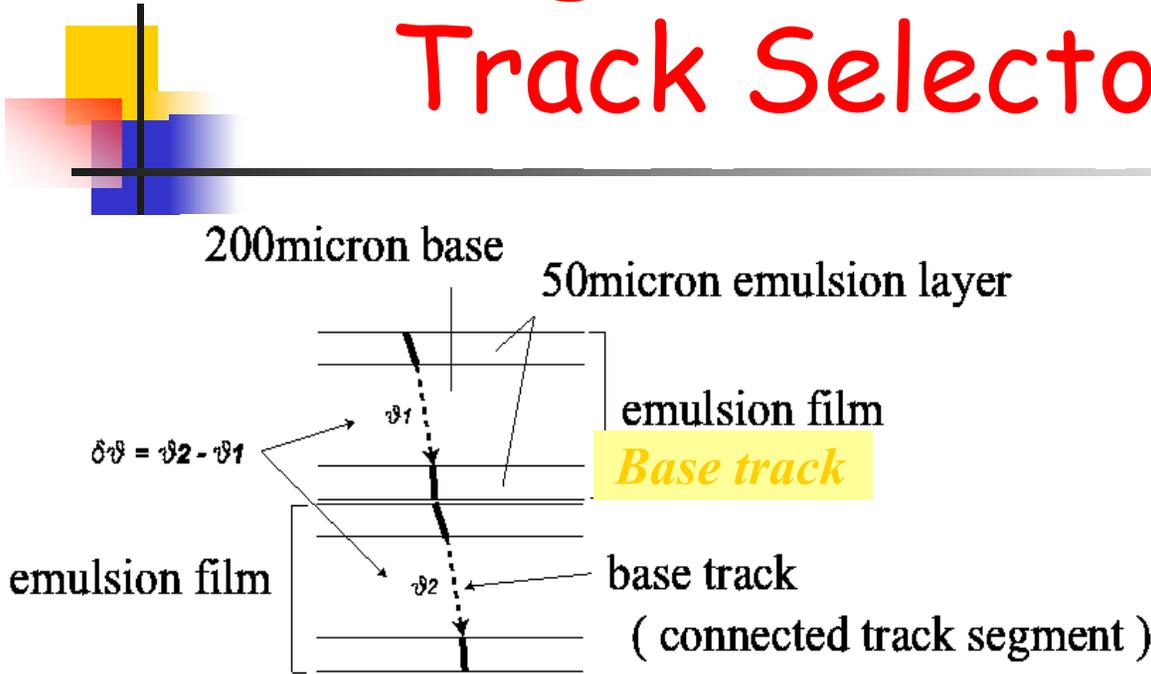
~ 30 grains / 100  $\mu\text{m}$   
grain diameter ~ 0.6  $\mu\text{m}$



**intrinsic tracking accuracy**

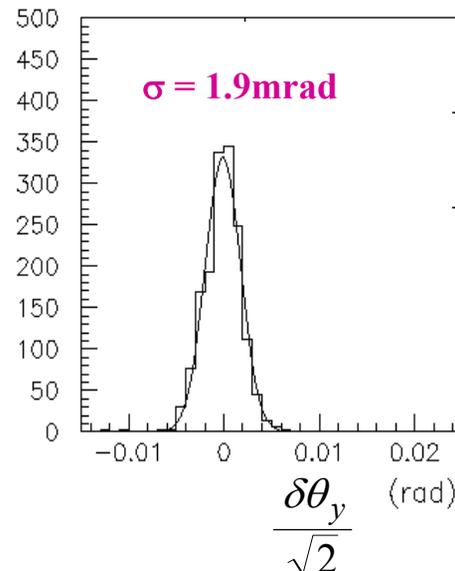
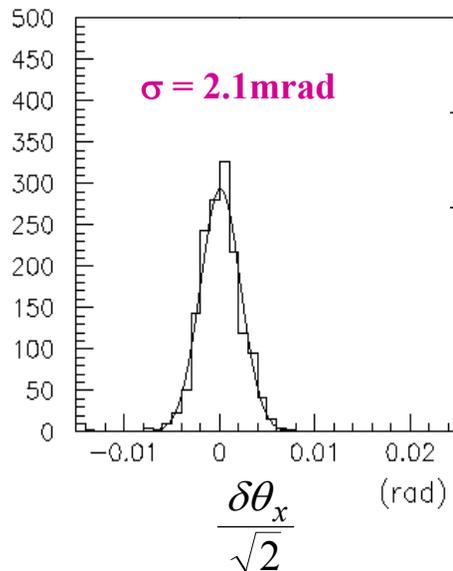
$$\sigma = 0.06 \mu\text{m}$$

# Angular resolution with Track Selector readout



Base track resolution  
(limited by digitisation error  
in image processing)

$$\sigma(\text{angle}) = 2.1 \text{ mrad}$$
$$\sigma(\text{position}) = 0.21 \mu\text{m}$$



Precision  
measurements  
for small samples



resolution  
close to the intrinsic  
resolution

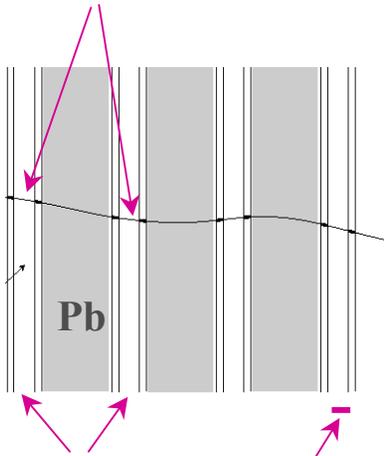
# Momentum measurement from Multiple Coulomb Scattering

(allows additional b.g. rejection by kinematic criteria)

## Angular method

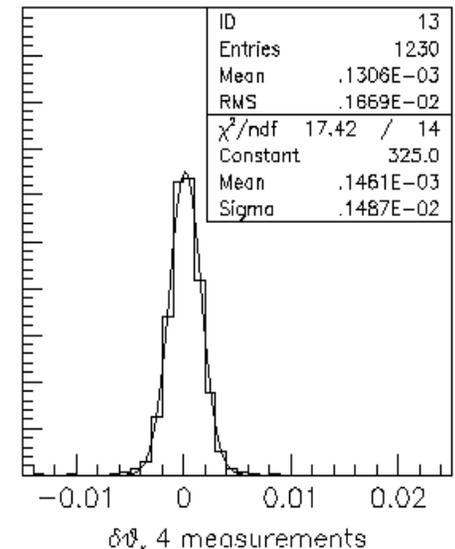
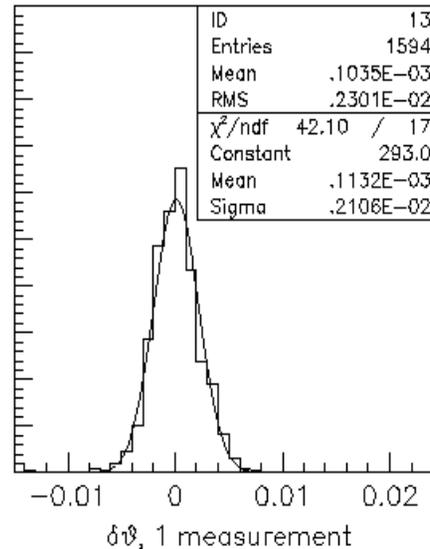
{ Not relying on alignment of emulsion films  
 Relying on parallelism of Pb plates and emulsion films

Angle difference



Film with emulsion layer on both sides of a plastic base

Plastic base thickness:  
lever arm for  
angle measurement

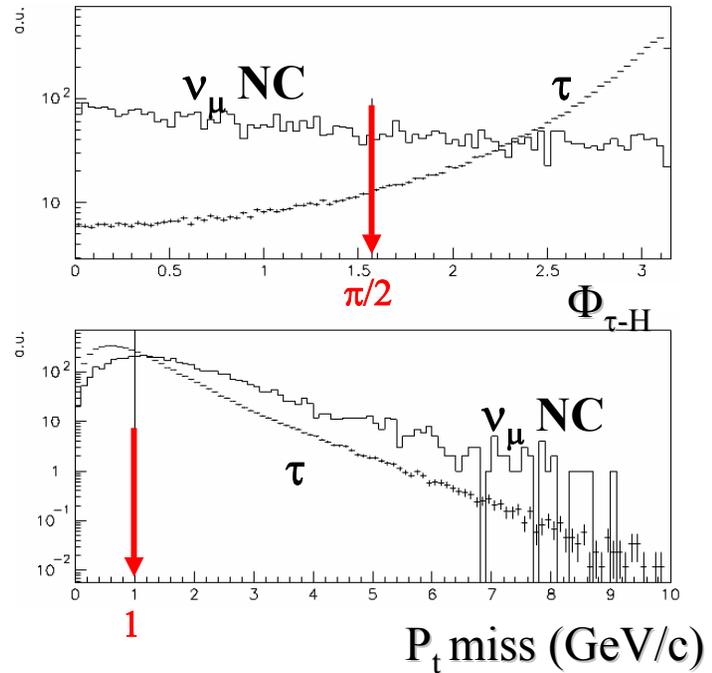
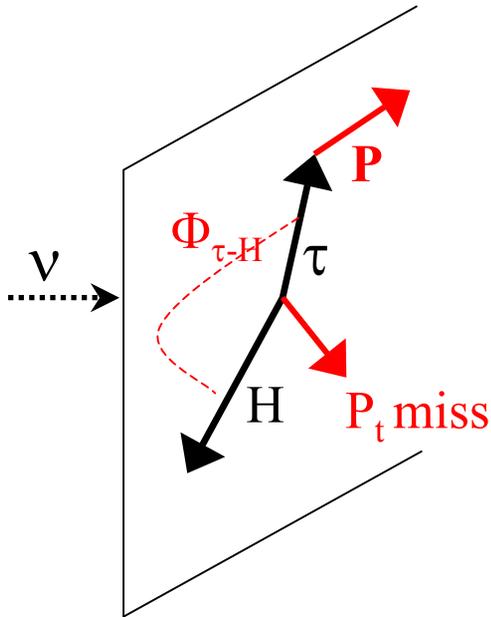


With intrinsic resolution 10.0 GeV/c

(max p measured with  $\Delta p/p < 0.2$  after  $5X_0$ )

# Global kinematics for $\tau \rightarrow h$

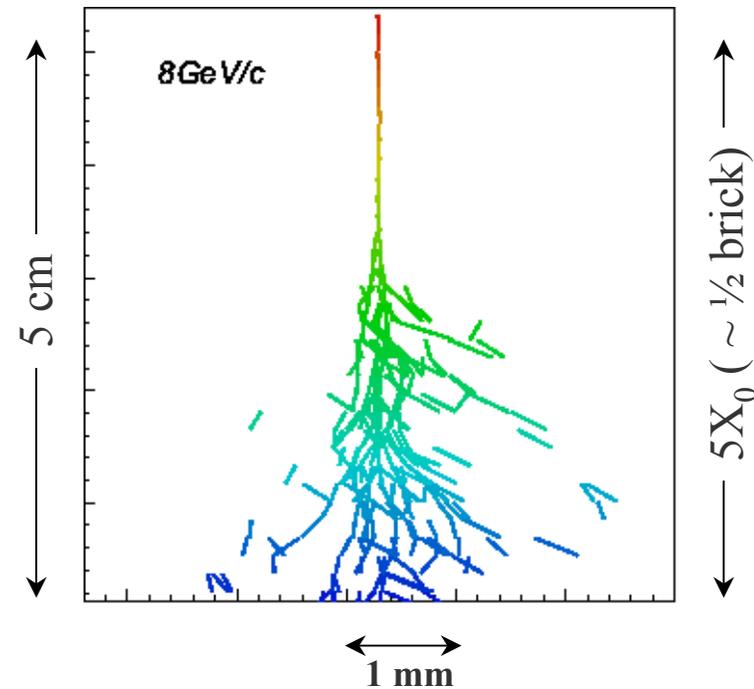
(for events with a  $\tau$  decay candidate)



This is just an example: it works also (event better) for  $\nu_e \text{ CC}$  interactions

# Electron identification and energy measurement (1)

- Electron identification close the Pb critical energy
- Performance estimated by reproducing the full chain:
  - Fuji-emulsions stored for about 2 months
  - Emulsion refreshing at the Nagoya University
  - Transportation of packed emulsions to CERN
  - Test exposure at the PS beam (mixture of  $e$  and  $\pi$ ,  $\checkmark$  upstream of the ECC)
  - Emulsion developed soon after the beam exposure
  - Scanning and analysis



# Electron identification and energy measurement (2)

## Identification

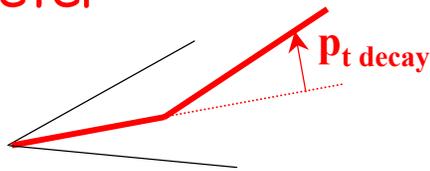
- Method based on shower identification and on MCS of the track
- $e/\pi$  ratio is measured with Cerenkov and ECC (good agreement)
  - ECC  $1.42 \pm 0.17$       Cerenkov  $1.46 \pm 0.11$       at 2GeV
  - ECC  $0.41 \pm 0.05$       Cerenkov  $0.32 \pm 0.03$       at 4GeV
- $\varepsilon_e = 88$  (91) % at 2 (4) GeV (in agreement with MC)

## Energy

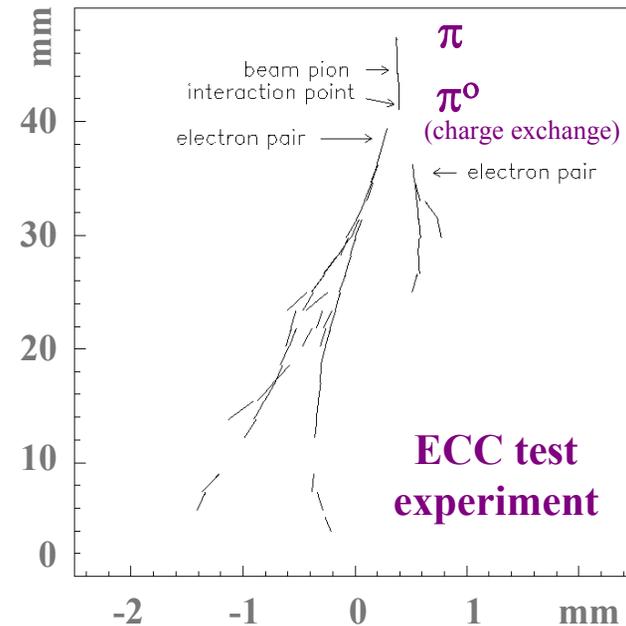
- Measured by counting the number of track segments into a cone along the electron track
- Multiple Coulomb Scattering before showering
- Resolution  $\sim 20\%$

# $\gamma/\pi^0$ reconstruction

- $\gamma$ s assigned to primary or to decay vertex depending on Impact Parameter



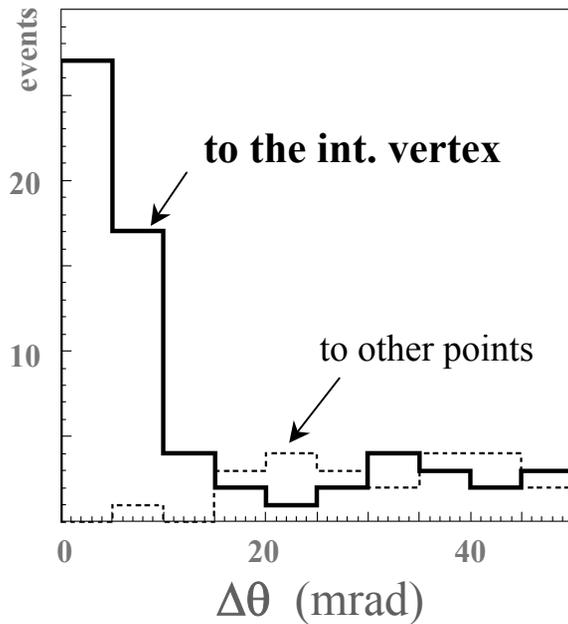
- $\gamma$ s assigned to the decay vertex
  - improved  $p_{t \text{ decay}}$  resolution (charged+neutral)
  - looser cut and higher efficiency
- $\gamma$ s assigned to the primary vertex
  - improved missing  $p_{\perp}$  resolution



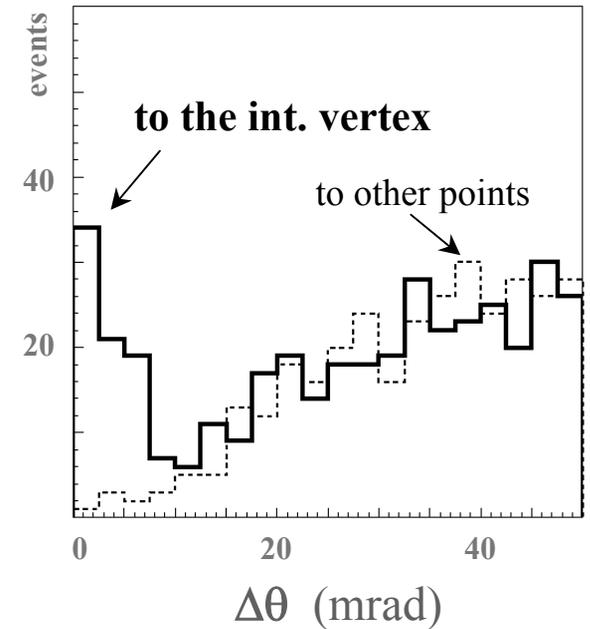
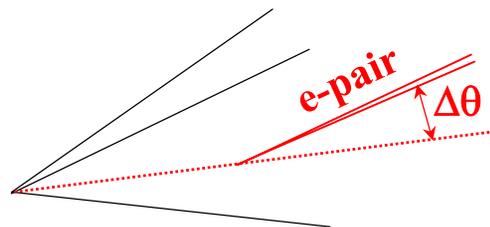
$\pi^0$  mass reconstruction also possible ( $\Delta m \sim 30\%$ ), but not used in the following analysis

# Pointing accuracy to the vertex of $e$ -pairs from $\gamma$ conversions

Studied in CHORUS and DONUT by NetScan  
( $\frac{1}{2} X_0$  depth in ECC)



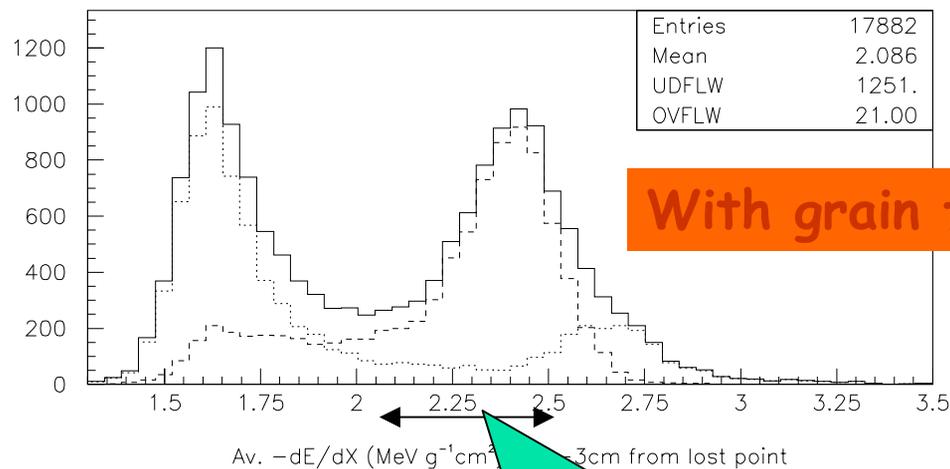
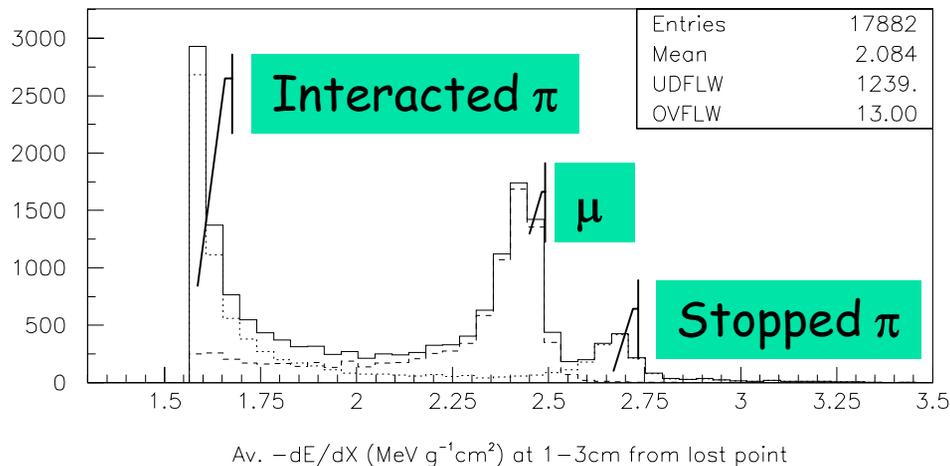
CHORUS

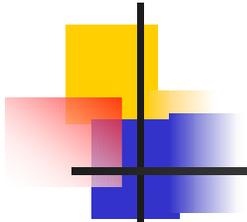


DONUT  
(ECC Fe-emulsion)

Important for increasing the sensitivity to  $\tau \rightarrow h n \pi^0$

# Averaged $dE/dX$ (under study)





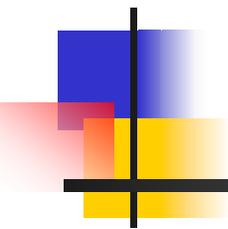
# Summary of the event reconstruction with OPERA

---

- High precision tracking ( $\delta x < 1\mu\text{m}$   $\delta\theta < 1\text{mrad}$ )
  - Kink decay topology
  - Electron and  $\gamma/\pi^0$  identification
- Energy measurement
  - Multiple Coulomb Scattering
  - Track counting (calorimetric measurement)
- Ionization ( $dE/dx$  measurement)
  - $\pi/\mu$  separation
  - $e/\pi^0$  separation

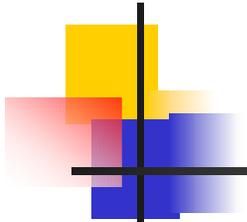


Topological and kinematical analysis event by event



$\nu_{\mu} \rightarrow \nu_e$  search with OPERA

---

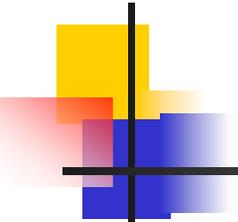


# Backgrounds

---

- $\pi^0$  identified as electrons produced in  $\nu_\mu$ NC and  $\nu_\mu$ CC with the  $\mu$  not identified
- $\nu_e$  beam contamination
- $\tau \rightarrow e$  from  $\nu_\mu \rightarrow \nu_\tau$  oscillations

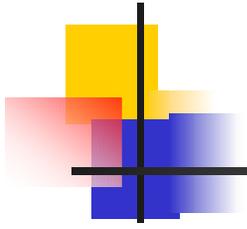
In the following we assume a three family mixing scenario with  $\theta_{23}=45^\circ$



# Background (1)

---

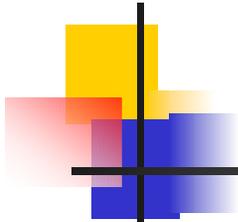
- Background from  $\pi^0$  in both  $\nu_\mu$ NC and  $\nu_\mu$ CC events classified as NC
  - $\gamma$ 's (from  $\pi^0$ ) converting to  $e^+e^-$  into the lead plate where the interaction occurred could be a large source of background.
  - The pairs can be identified
    - if the opening angle ( $\Delta\theta$ ) is larger than 3mrad. This cut is fixed by the angular resolution achievable with emulsion films
    - if  $\Delta\theta$  is smaller than 3mrad, we exploit the emulsion capability to count grains associated to a track. Having 30grains/100 $\mu$ m we require for a single electron that  $\#grains/100\mu m < 30 + 3 \times 30^{\frac{1}{2}}$ . Conservatively we use only the emulsion film downstream from the vertex lead plate
  - A cut  $E_e > 1$  GeV is also applied to reduce the soft  $\gamma$  component
  - A cut on the missing  $p_T$  ( $< 1.5$  GeV) is also applied to further reduce the NC component



# Backgrounds (2)

---

- $\nu_e$  CC from the beam contamination
  - Very difficult to suppress. Identical to the signal we are looking for!
  - CC interactions of  $\nu_e$  coming from oscillations are softer
    - $E_{vis}$  has to be smaller than 20 GeV
- $\tau \rightarrow e$  from  $\nu_\mu \rightarrow \nu_\tau$  oscillations
  - its amount depends on  $\Delta m^2_{23}$ . In the following we assume  $\Delta m^2_{23} = 2.5 \times 10^{-3} \text{ eV}^2$
  - in OPERA it is largely suppressed because only events with the decay kink not identified contribute to the background
  - it is slightly reduced by the missing  $p_T$  cut



# $\nu_\mu \rightarrow \nu_e$ : selection efficiencies

Location eff.

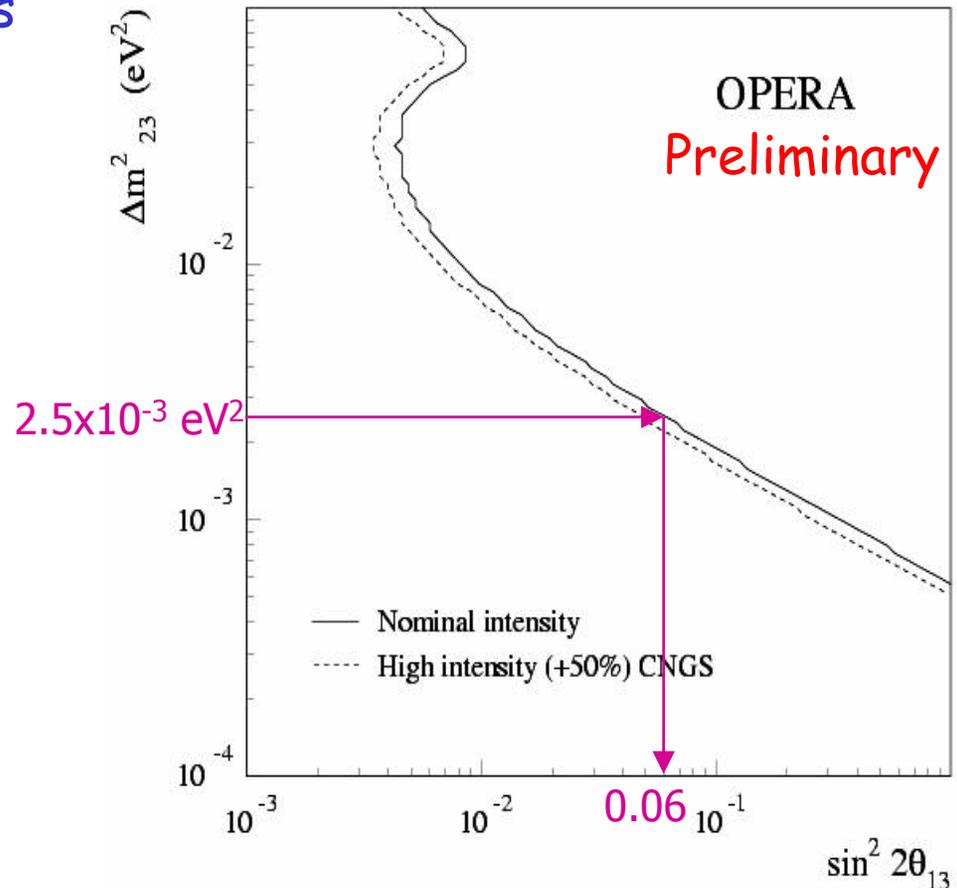
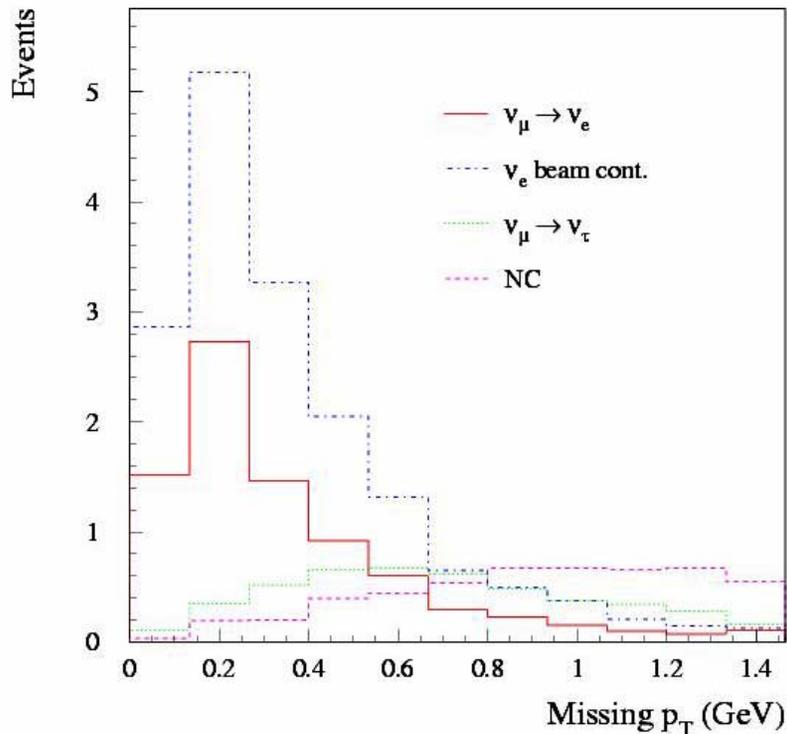
	signal	$\tau \rightarrow e$	$\nu_\mu$ CC	$\nu_\mu$ NC	$\nu_e$ CC beam
$\xi$	0.53	0.053	0.52	0.48	0.53
Total eff.	0.31	0.032	$0.34 \times 10^{-4}$	$7.0 \times 10^{-4}$	0.082

Expected signal and background assuming 5 years data taking with the nominal CNGS beam and  $\Delta m^2_{23} = 2.5 \times 10^{-3} \text{ eV}^2$ ,  $\sin^2 2\theta_{23} = 1$

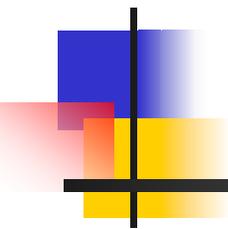
$\theta_{13}$	signal	$\tau \rightarrow e$	$\nu_\mu$ CC	$\nu_\mu$ NC	$\nu_e$ CC beam
$9^\circ$	9.3	4.5	1.0	5.2	18
$8^\circ$	7.4	4.5	1.0	5.2	18
$7^\circ$	5.8	4.6	1.0	5.2	18
$5^\circ$	3.0	4.6	1.0	5.2	18
$3^\circ$	1.2	4.7	1.0	5.2	18

# OPERA sensitivity to $\theta_{13}$

By fitting simultaneously the  $E_e$ , missing  $p_T$  and  $E_{vis}$  distributions we got the sensitivity at 90%

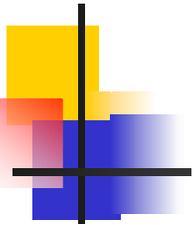


$\nu_{\mu} \rightarrow \nu_e$  search combining  
ICARUS and OPERA



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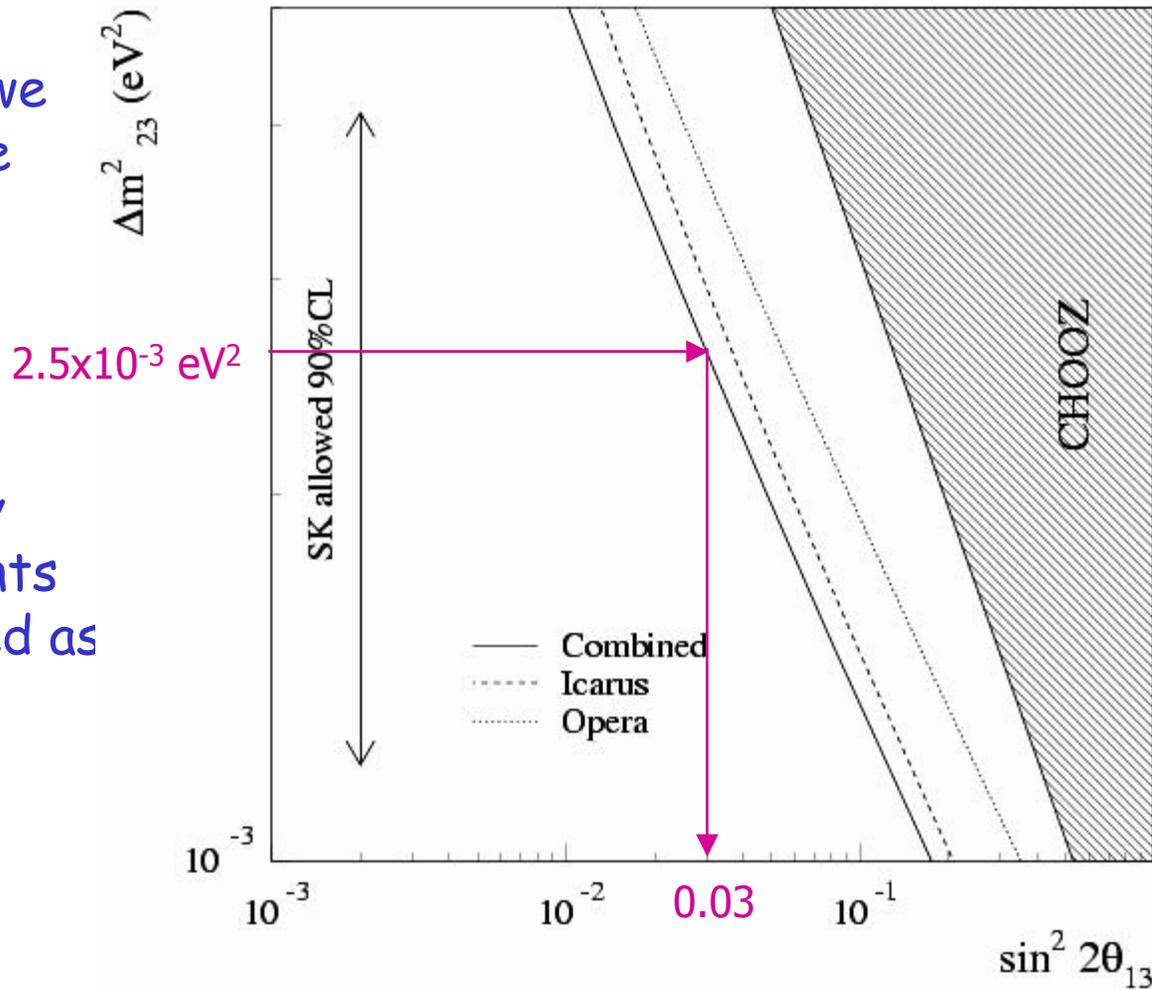
# ICARUS and OPERA combined sensitivity



CNGS 5 years nominal beam

- By using the information reported in the proposal we managed to reproduce the ICARUS sensitivity curve at 90% C.L.
- Finally, we evaluated the CNGS sensitivity to  $\theta_{13}$  by combining both experiments through a global  $\chi^2$  defined as

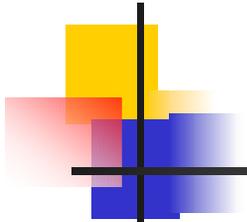
$$\chi^2 = \chi^2_{\text{ICARUS}} + \chi^2_{\text{OPERA}}$$



# Comparing different scenarios

Limits at 90% C.L. on  $\sin^2 2\theta_{13}$  and  $\theta_{13}$   
( $\Delta m^2_{23} = 2.5 \times 10^{-3} \text{ eV}^2$ ;  $\sin^2 \theta_{23} = 1$ )

Experiment	$\sin^2 2\theta_{13}$	$\theta_{13}$
CHOOZ	$<0.14$	$<11^\circ$
MINOS 2yr	$<0.06$	$<7.1^\circ$
ICARUS 5yr	$<0.04$	$<5.8^\circ$
OPERA 5yr	$<0.06$	$<7.1^\circ$
ICARUS+OPERA 5yr $\equiv 3\text{yr CNGS} \times 1.5$	$<0.03$	$<5.0^\circ$
ICARUS+OPERA 5yr CNGS $\times 1.5$	$<0.025$	$<4.5^\circ$
ICARUS 5yr Low energy CNGS	$<0.02$	$<4.1^\circ$
JHF 5yr	$<0.006$	$<2.5^\circ$

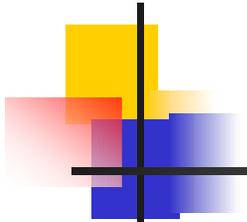


# Time is an important issue

Sensitivity to  $\theta_{13}$  as a function of the year

	2004	2005	2006	2007	2008	2009
CHOOZ	<0.14	<0.14	<0.14	<0.14	<0.14	<0.14
MINOS		→	<0.085	<0.06	<0.049	<0.042
CNGS*			→	<0.067	<0.047	<0.039
CNGSx1.5*			→	<0.056	<0.039	<0.033
Low energy CNGS				→	<0.040	<0.028
JHF					→	<0.013

\* Designed for  $\nu_{\tau}$  appearance



# Conclusion

---

- OPERA has sensitivity to  $\theta_{13}$  comparable to the one of the other LBL experiments
  - $\sin^2 2\theta_{13} < 0.05-0.06$
- The sensitivity to  $\theta_{13}$  combining ICARUS and OPERA is
  - $\sin^2 2\theta_{13} < 0.03$
- If the CNGS starts in 2006, by the time JHF will provide the first results (2009) the combined sensitivity of ICARUS and OPERA at  $\Delta m^2_{23} = 2.5 \times 10^{-3} \text{ eV}^2$ ,  $\sin^2 \theta_{23} = 1$  will be a factor 4-5 better than the CHOOZ limit
- If we are lucky ( $\theta_{13}$  just below the CHOOZ limit) the CNGS could provide the first measurement of  $\theta_{13}$ !